Chirality dependence of the *K***-momentum dark excitons in carbon nanotubes**

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Using a collection of 12 semiconducting carbon-nanotube samples, each highly enriched in a single chirality, we study the chirality dependence of the *K*-momentum dark singlet exciton using phonon sideband optical spectroscopy. Measurements of bright absorptive and emissive sidebands of this finite momentum exciton identify its energy as 20–38 meV above the bright singlet exciton, a separation that exhibits systematic dependencies on tube diameter, 2*n*+*m* family, and chiral index. We present calculations that explain how chiral angle dependence in this energy separation relates to the Coulomb exchange interaction and elaborate the dominance of the $K_{A_1'}$ phonon sidebands over the zone-center phonon sidebands over a wide range of chiralities. The Kataura plot arising from these data is qualitatively well described by theory but the energy separation between the sidebands shows a larger chiral dependence than predicted. This latter observation may indicate a larger dispersion for the associated phonon near the *K* point than expected from finite distance force modeling.

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I. INTRODUCTION

Semiconducting carbon nanotubes (CNTs) have attracted widespread interest due to their unique electronic structure. Their quasi-one-dimensional (1D) nature makes electronelectron interactions important, leading to the formation of strongly bound exciton states. The absorption and emission spectra of CNTs are well known to be dominated by excitons, allowing for a direct study of many-body physics in a quasi-1D regime. $1-3$ The lowest energy manifold of excitons is comprised of 16 states: four singlets and 12 triplets. Only one of these excitons is bright (optically active). $4-6$ $4-6$ The remaining 15 excitons are dark (optically inactive) and have remained largely unstudied for this reason.

Experimental investigations of dark excitons initially focused on the zero-momentum dark singlet exciton, which can be brightened by the application of an external magnetic field[.7–](#page-8-5)[9](#page-8-6) Lattice defects induced by intense pulsed laser exci-tation can also brighten otherwise dark exciton states.^{10[,11](#page-8-8)} Beyond the zero-momentum dark singlet, there remain two time reversal degenerate dark singlet excitons with center-ofmass momenta near the K and K' points of the graphene Brillouin zone (BZ). While these *K*-momentum dark singlet excitons (at energy E_K) cannot be directly photoexcited due to momentum conservation, they can be excited in combination with a *K*-momentum phonon. This process is enhanced by the strong exciton-phonon coupling near the *K* points of the graphene BZ, resulting in the appearance of pairs of bright phonon sidebands located *asymmetrically* about the bright exciton.^{12[,13](#page-8-10)} Previously, several of us used these sideband pairs to study the *K*-momentum exciton in a purified $(n,m) = (6,5)$ CNT sample, where (n,m) denotes the CNT wrapping vector.¹³ The amplitude, shape, and separation of the sidebands were found to be in good agreement with a theory modeling them as bright sidebands of the dark *K*-momentum singlet exciton. Nevertheless, the inferred separation of E_K from the bright singlet (E_{11}) exceeded theoretical predictions and the sidebands have had other inter-pretations in the literature.^{14–[18](#page-8-12)} Hence, a more comprehensive study over a wider range of chiralities is called for.

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Figure [1](#page-0-0) illustrates our approach in using sideband spectroscopy to identify E_K . Phonons of A'_1 symmetry at the K' troscopy to identify E_K . Phonons of A'_1 symmetry at the K'
 (K) point (iTO branch) create dynamic $\sqrt{3} \times \sqrt{3}$ intervalley scattering potentials that couple excitons at the $K(K')$ point to excitons at the zone center. For a visualization of this phonon mode, the reader is referred to Fig. 7 of Ref. [19,](#page-8-13) which illustrates the atomic displacements labeled as K_1 .) The associated hybridization creates bright phonon sidebands symmetrically about E_K . We denote the lower-energy emissive sideband X_1 , and the upper-energy absorptive sideband X_2 . One can think of X_2 as creating a phonon-exciton pair, each at the zone boundary but with zero combined momentum, a transition made possible through their mixing with the bright exciton. Similarly, X_1 corresponds to an emissive process in which the *K*-momentum exciton creates a photon but conserves momentum by emitting a phonon. As indicated in the level diagram, the energetic separation between X_1 and X_2 is essentially twice the $K_{A'_1}$ phonon energy, and measurement of both sidebands permits an interpolation

FIG. 1. (Color online) Energy-level diagram of the CNT excitonic states and phonon sidebands. The solid double arrow indicates optical transitions involving the bright exciton. The dotted arrows denote absorptive and emissive transitions involving the sidebands X_1 and X_2 of the dark exciton at E_K . $\hbar \omega_K$ is the energy of a $K_{A_1'}$ phonon.

of E_K . Two recent works have studied one of these sidebands, X_1 , over a range of chiralities but did not reveal a systematic chirality dependence.^{11,[20](#page-8-14)} In contrast, here sideband *pairs* are simultaneously identified in 12 highly purified single-chirality CNT suspensions permitting the determination of E_K . Our high-resolution study finds that both E_{X_1} and E_{X_2} depend on chirality, and that qualitative agreement with theory emerges in the pattern of energies produced by a comprehensive study of E_K . These findings lead to a physical understanding of the processes that dominate this dark-bright splitting. Specifically, whereas the chiral-family behavior of the bright exciton emerges predominantly from the single quasiparticle energies, the dark-bright splitting measured here reveals chiral behavior arising from the many-particle exchange interaction.

Our study shows that X_1 and X_2 appear asymmetrically about the bright exciton for each species, and exhibit systematic chiral dependences. Kataura plots of $\Delta_K = E_K - E_{11}$ show $2n+m$ family behavior, chiral-index dependence, and an overall scaling with diameter. A recent theoretical study²¹ and our own extended tight-binding (ETB) calculations predict behavior consistent with these results, and a $1/d^2$ scaling of Δ_K in the large diameter limit. Because the X_2 absorptive sideband has been interpreted as a $\Gamma_{E_{2g}}$ (longitudinal optical mode at the Γ -point) phonon sideband of the bright exciton, $14-16$ we calculate the relative contributions of the $\Gamma_{E_{2g}}$ and $K_{A'_1}$ phonons to X_2 over a large diameter range, and find the $K_{A_1'}$ sideband dominates by more than an order of magnitude in all cases. This prediction, along with the rather strong chiral dependence of the experimental X_2 sideband energy, strongly discourages against further interpretation of X_2 as arising from zone-center phonons. The energetic spacing between X_1 and X_2 should be given by two $K_{A'_1}$ phonon energies. Experimentally, we note that this separation shows a larger chiral dependence than expected, suggesting the $K_{A_1'}$ phonons may be more dispersive near the *K* point than current finite distance force models predict.

II. EXPERIMENTAL METHODS

Highly purified single-chirality dispersions of $(6,5)$, $(7,5)$, $(7,6)$, $(8,3)$, $(8,4)$, $(8,6)$, $(8,7)$, $(9,1)$, $(9,4)$, $(9,5)$, $(10,2)$, and (10,5) CNTs are prepared following the method of Ref. [22.](#page-8-17) The output of a 150 W Xenon arc lamp passes through the first two 0.5-m stages of a triple spectrometer in a bandpass configuration. The resulting 10-nm-wide excitation spectrum contains very low spectral content outside the bandpass window and is further filtered to remove second-order diffractive components. Aqueous suspensions are excited in precision quartz cuvettes with excitation and emission beams at 90°. Photoluminescence (PL) is collected into the third stage of our triple spectrometer and detected with a TE-cooled silicon charge coupled device (visible) and a Li N_2 -cooled InGaAs array (near infrared). An additional 0.5-m spectrometer and TE-cooled InGaAs array are used as a power meter. To ensure proper averaging over the spatial profile of the excitation beam, a beam-splitter samples part of the excitation and focuses it onto a diffuser first before collection into the

FIG. 2. (Color online) PL maps of (a) $(9,4)$, (b) $(10,2)$, (c) $(8,6)$, and (d) (8,4) CNTs. The excitation line runs from upper right to bottom left and the E_{11} peak of the dominant species shows as the brightest spot along the excitation line. Dashed lines indicate energies in eV, as labeled.

power meter. All elements of our setup are calibrated for spectral response with a NIST traceable light source and all reported spectra are power normalized. The *E*¹¹ absorption resonance is frequently used in our analysis and is typically determined using PL excitation profiles discussed in Sec. [III.](#page-1-0) We find that optical absorption spectra give the same value of *E*¹¹ and are used as a supplemental measurement. Absorption data collected for this purpose are obtained using a tungsten lamp and the aforementioned triple 0.5-m spectrometer and detectors.

By combining PL emission spectra taken at different excitation energies, we assemble two-dimensional excitationemission PL maps. In our study the range of excitation energies usually passes directly through the *E*¹¹ resonance, known as resonant PL spectroscopy. Typically, broadband excitation scatter would limit the efficacy of resonant PL spectroscopy for these samples by obscuring interesting features close to the excitation energy. Here, the low spectral background of our excitation source makes resonant PL spectroscopy an effective tool for studying X_1 and X_2 X_2 [Fig. 2].

III. DATA ANALYSIS

Theoretical calculation methods described in Refs. [12,](#page-8-9) [23,](#page-8-18) and [24](#page-8-19) and adopted by several of us in Ref. [13](#page-8-10) are frequently used in our analysis. We elaborate these methods here for completeness. First, we obtain a single-particle basis of states using an ETB model with a nearest-neighbor overlap integral of 2.7 eV. We then use the Bethe-Salpeter equation to calculate exciton energies E_q^S and wave functions

$$
|\Psi_{q}^{S}\rangle = \sum_{\mathbf{k}} A_{\mathbf{k},\mathbf{q}}^{S} c_{\mathbf{k}+\mathbf{q},c}^{\dagger} c_{\mathbf{k},v} |\mathbf{G}\mathbf{S}\rangle \tag{1}
$$

following the methods of Ref. [24.](#page-8-19) Here, $A_{\mathbf{k},\mathbf{q}}^{S}$ are the exciton wave-function coefficients for and electron-hole pair with center-of-mass momentum **q** and relative momenta **k**+**q**/**2**, *S*=0,1,... indexes the ground and excited states of the exciton, and $c_{\mathbf{k},i}^{\dagger}$ is an electron creation operator for wave vector **k** in the conduction $(i=c)$ or valence $(i=v)$ band. Our simulation uses 20 000 atoms and an environmental dielectric constant of $2.7¹³$ We compute the phonon eigenspectrum using a symmetry-adapted valence force constant model²³ up to fifth nearest neighbors.¹⁹ We then calculate the electronphonon couplings $M_{\mathbf{k},\mathbf{q}}^{\mu}$ to first order in the lattice deformations[,25](#page-8-20) with a nearest-neighbor coupling constant of 5.3 eV/ \AA .^{[12](#page-8-9)} Here, μ =1,2,...,6 is the phonon band index. The electron-phonon coupling is then used to compute the *exciton*-phonon coupling¹²

$$
B_{\mathbf{q}\mu}^{SS'} = \sum_{\mathbf{k}} M_{\mathbf{k},\mathbf{q}}^{\mu} A_{\mathbf{k},\mathbf{q}}^{S'*}(A_{\mathbf{k},0}^S + A_{\mathbf{k}+\mathbf{q},0}^S). \tag{2}
$$

As described in Ref. [12,](#page-8-9) mixing of the bright exciton with excitons at finite momentum **q** and phonons at −**q** gives an unbroadened absorptive sideband line shape of the form

$$
I(\omega) \propto \delta(E_{11} - \hbar \omega) + \sum_{\mathbf{q}, \mu, S'} \frac{|B_{\mathbf{q}, \mu}^{0S'}|^2}{(E_{\mathbf{q}}^{S'} + \hbar \omega_{-\mathbf{q}}^{\mu} - E_{11})^2}
$$

$$
\times \delta(E_{\mathbf{q}}^{S'} + \hbar \omega_{-\mathbf{q}}^{\mu} - \hbar \omega). \tag{3}
$$

In the above expression, $\omega_{\text{-}q}^{\mu}$ is the phonon eigenfrequency at wave vector $-\mathbf{q}$ and E_{11} is the bright exciton energy.¹² The second term represents the X_2 sideband line shape and in modeling our data we convolve this form with a Gaussian to represent inhomogeneous broadening due to potential fluctuations. Our calculation also produces theoretical values of energies E_{X_2} , E_K , and E_{11} .

We measure the energy of the *K*-momentum exciton through phonon sideband optical spectroscopy, in which the average energy of X_1 and X_2 gives E_K , and these features are both separated from E_K by the energy of one $K_{A'_1}$ phonon.¹³ Representative PL maps, resonant PL spectra, and photoluminescence excitation (PLE) spectra from our measurements are shown in Figs. [2–](#page-1-1)[4,](#page-2-0) respectively. PL maps taken on all samples show X_1 and X_2 features resonantly aligned with E_{11} absorption or emission, respectively, associating these transitions with the dominant chirality of the sample. A subset of these maps appears in Fig. [2](#page-1-1) with X_1 and X_2 identified by arrows in Fig. $2(a)$ $2(a)$. The energetic location of these sidebands relative to E_{11} is always asymmetric, in accord with Ref. [13](#page-8-10) and placing E_K consistently above E_{11} . Overlapping transitions from unintended chiralities are suppressed by using samples highly enriched in a single chirality. Further discrimination against features associated with other chiralities is obtained by analyzing vertical or horizontal line cuts through the data that pass through the E_{11} resonance [Fig.

FIG. 3. On-resonance PL spectra for (a) $(7,5)$, (b) $(9,5)$ (c) $(8,3)$, and (d) (10,5) CNTs.

 $2(c)$ $2(c)$]. Emission (excitation) spectra containing the X_1 (X_2) sideband are thereby obtained at the excitation (emission) energy that places E_{11} and X_1 (X_2) into resonance, a subset of which are shown in Fig. 3 (Fig. [4](#page-2-0)). The location of these line cuts is indicated in Fig. $2(c)$ $2(c)$, with the spectrum resulting from a X_1 cut heretofore referred to as an "on-resonance" PL spectrum, and that arising from a X_2 cut denoted a PLE spectrum. PLE is essentially an absorbance spectroscopy but gives superior background rejection compared to conventional absorbance as the latter cannot discriminate among contributions from different chiralities.

Note that here we are interested in the relative energy separation of X_1 and X_2 but X_1 contains a Stokes shift since emission resonances in CNTs generally shift a few meV

FIG. 4. PLE spectra for (a) $(8,3)$, (b) $(8,6)$ (c) $(8,4)$, and (d) $(9,4)$ CNTs.

FIG. 5. (Color online) PL (black, dashed) and PLE (red, solid) spectra of $(8,6)$ CNTs plotted on an energy axis relative to E_{11} . Vertical lines label E_{X_1} , E_{X_2} , E_{11} , and E_K . The blue dashed doublearrowed lines label the energetic spacing of the sidebands from E_K , and $\hbar \omega_K$ is the energy of a $K_{A_1'}$ phonon. Note that the peak of X_2 is offset from E_{X_2} as discussed in the text.

lower in energy than their absorptive counterparts. A relative energy scheme, as shown in Fig. [5,](#page-3-0) then becomes quite useful. In this plot, X_1 and X_2 are placed on a common energy axis and measured relative to $E_{11}^{(PL)}$ and $E_{11}^{(PLE)}$, respectively.¹³ Here, $E_{11}^{(PL)}$ ($E_{11}^{(PLE)}$) is the energetic location of E_{11} that ap-pears in the PL (PLE) spectra and is summarized in Table [I.](#page-3-1) To avoid problems discriminating PL emission from excitation scatter at the peak of the E_{11} resonance, we actually determine $E_{11}^{(PL)}$ and $E_{11}^{(PLE)}$ by taking line cuts across the relatively weaker X_2 and X_1 features at locations indicated in Fig. $2(b)$ $2(b)$ and performing Gaussian or Lorentzian fits. We have found that absorbance provides the same values of $E_{11}^{(PLE)}$ as these line cuts and because chiral discrimination is not an issue in determining the location of the strong E_{11} feature, we use whichever method gives the least error in determining $E_{11}^{\text{(PLE)}}$.

We remark that the PLE spectra shown in Figs. [4](#page-2-0) and [5](#page-3-0) all show evidence of a small feature \sim 90 meV above $E_{11}^{(PLE)}$.

Our theoretical calculations predict the presence of a phonon sideband of E_K at this energy, arising from *K*-momentum oTA and oTO phonon modes. However, our modeling suggests the intensity of this feature relative to X_2 is only \sim 1/200 and would be unobservable in these measurements. Hence, the identification of this minor sideband remains an open question to the community.

Some analysis is required to accurately determine energetic locations of X_1 and X_2 , denoted E_{X_1} and E_{X_2} , respectively. Theory suggests these features should lie symmetrically about E_K but their line shapes differ significantly. X_2 is an absorptive process with a continuum of states available above a lower-energy threshold and the resulting asymmetric line shape follows closely the density of available excitonphonon final states.¹³ X_1 is an emissive process so that thermal relaxation drives the excited population toward the lowest-energy excitons having center-of-mass momenta closest to the *K* point. The resulting more symmetric line shape thus resembles that of E_{11} . These spectral profiles influence our determination of E_{X_1} and E_{X_2} . In the case of E_{X_1} , X_1 happens to lie on the tail of the E_{11} emission. We thus subtract a linear background before fitting X_1 with a Lorentzian to obtain its central value.

The determination of E_{X_2} includes an offset that is related to line broadening of an asymmetric line shape. To understand this relationship, imagine an asymmetric unbroadened profile resembling $\Theta(E-E_1) \exp(-E/E_0)$, where Θ is a step function. If we apply a Gaussian broadening to this function the resulting line shape will peak above E_1 . The profile of X_2 before broadening is qualitatively similar to this function, with an absorbance threshold at $E_{X_2} = E_K + \hbar \omega_K$, where $\hbar \omega_K$ is the energy of the $K_{A_1'}$ phonon closest to the *K* point. We find that 30-meV Gaussian broadening produces a \sim 8 meV blueshift of the peak of X_2 relative to E_{X_2} , which is rather insensitive to chirality. A broadening of 42 meV gives a larger 11-meV shift and is used here because it produces the best fit to X_2 [Fig. [6](#page-4-0)]. To determine E_{X_2} , we fit X_2 to determine the location of its peak and then subtract the aforementioned shift, explicitly computed for that particular chirality. The X_2 fitting line shape is used only to determine the peak position, is generic to all chiralities, and is calculated for a

(n,m)	$E_{11}^{\rm (PLE)}$	$E_{11}^{\rm (PL)}$	E_{X_1}	E_{X_2}	Δ_K
(6,5)	1.258 ± 0.0000	1.253 ± 0.0001	1.118 ± 0.0002	1.451 ± 0.0034	0.0291 ± 0.0034
(7,5)	1.200 ± 0.0002	1.197 ± 0.0001	1.063 ± 0.0013	1.400 ± 0.0026	0.0329 ± 0.0029
(7,6)	1.099 ± 0.0004	1.095 ± 0.0001	0.960 ± 0.0008	1.281 ± 0.0016	0.0234 ± 0.0018
(8,3)	1.286 ± 0.0001	1.283 ± 0.0001	1.151 ± 0.0004	1.487 ± 0.0020	0.0342 ± 0.0021
(8,4)	1.110 ± 0.0001	1.105 ± 0.0001	0.968 ± 0.0004	1.291 ± 0.0012	0.0222 ± 0.0013
(8,6)	1.054 ± 0.0001	1.049 ± 0.0002	0.916 ± 0.0010	1.244 ± 0.0013	0.0280 ± 0.0017
(8,7)	0.970 ± 0.0001	0.969 ± 0.0002	0.834 ± 0.0005	1.154 ± 0.0013	0.0247 ± 0.0014
(9,1)	1.343 ± 0.0005	1.340 ± 0.0001	1.208 ± 0.0007	1.546 ± 0.0019	0.0355 ± 0.0021
(9,4)	1.115 ± 0.0002	1.111 ± 0.0003	0.978 ± 0.0009	1.307 ± 0.0026	0.0295 ± 0.0028
(9,5)	0.989 ± 0.0000	0.987 ± 0.0001	0.852 ± 0.0007	1.174 ± 0.0029	0.0252 ± 0.0029
(10,2)	1.167 ± 0.0003	1.162 ± 0.0000	1.029 ± 0.0009	1.366 ± 0.0026	0.0331 ± 0.0028
(10,5)	0.983 ± 0.0000	0.982 ± 0.0002	0.846 ± 0.0013	1.168 ± 0.0018	0.0244 ± 0.0023

TABLE I. Measured values of $E_{11}^{(PLE)}$, $E_{11}^{(PL)}$, E_{X_1} , E_{X_2} , and Δ_K in eV.

FIG. 6. (Color online) Best fit of X_2 for $(8,6)$ CNTs with the theoretical X_2 lineshape described in the text.

(6,5) CNT. Ultimately, these data are used to locate the position of E_K and in that regard we emphasize two important points. First, variations in the shift with chirality are small and even if we do not account for it the family behavior of E_K remains the same. Second, systematic errors arising from the choice of broadening result in a shift of the inferred value of E_K that is constant across all chiralities and of order 1 meV, again leaving all reported family behavior unchanged.

Defining the splitting energy of the *K*-momentum exciton relative to the bright exciton as $\Delta_K = E_K - E_{11}$, we obtain $\Delta_{K} = (E_{X_1} + E_{X_2} - 2E_{11})/2$ (Fig. [5](#page-3-0)). Table [I](#page-3-1) summarizes $E_{11}^{(PL)}$, $E_{11}^{(PLE)}$, E_{X_1} , E_{X_2} , Δ_K , and their nonsystematic errors. Data for $(6,5)$ CNTs was taken from Ref. [13](#page-8-10) and reanalyzed following the above procedure. This results in a somewhat different, but more accurate, value of Δ_K than previously reported.¹³

IV. DISCUSSION AND SUMMARY

Experimental results for Δ_K versus inverse tube diameter $(1/d)$ are shown in Fig. [7](#page-4-1)(a). Predictions for Δ_K by an *ab initio* parameterized variational calculation²¹ and our ETB model are shown in Figs. $7(b)$ $7(b)$ and $7(c)$. In the experimental data, we observe a grouping of Δ_K by $2n+m$ families and a dependence upon chiral index (a so-called "Kataura plot"). The chiral index $\nu = (n - m)$ mod 3 typically classifies two distinct groups of semiconducting family behaviors. In our case $\nu = 1(2)$ -type semiconducting tubes have a Δ_K that fans downward (upward) in energy. The theoretical results of Ref. [21](#page-8-15) and our ETB calculations show the same semiconducting and family groupings seen in the data. The ETB model underestimates the value of Δ_K while the variational calculation is in much better agreement.²¹

 Δ_K is the difference between the binding energies of E_K and E_{11} , and thus a direct measure of electron-electron interactions in semiconducting CNTs. The physics governing many-particle interactions is contained within the interaction kernel, $K=2K^x-K^d$, of the Bethe-Salpeter equation.^{24[,26](#page-8-21)} $\mathcal{K}^d(\mathcal{K}^x)$ is the direct (exchange) interaction. In Appendix, we study the properties of the interaction kernel for E_{11} and E_K . The direct interaction is usually responsible for exciton

FIG. 7. (Color online) Δ_K versus $1/d$ for (a) experimental data, (b) *ab initio* predictions from Ref. [21,](#page-8-15) and (c) our ETB calculation. Lines connect families sharing the same value of $2n+m$ indicated in (a) .

 $1/d$ (nm⁻¹)

formation²⁶ except in unusual circumstances, 27 and has only a weak dependence on chiral angle for E_{11} and E_K . Indeed the direct energy is nearly identical in the two excitons and thus has little effect on Δ_K . The exchange interaction, however, differs significantly between the two excitons. For E_K , the exchange interaction involves elements of the Coulomb interaction in the vicinity of the *K* point. This makes the

FIG. 8. (Color online) Calculated exchange energy of (a) E_K and (b) E_{11} versus $1/d$. Lines connect families sharing the same value of $2n+m$ indicated in (a).

associated exchange energy [Fig. $8(a)$ $8(a)$] relatively sensitive to the threefold lattice anisotropy.²⁸ In contrast, the exchange interaction for E_{11} involves only small momentum properties of the interaction kernel, which are less sensitive to chiral angle variations [Fig. $8(b)$ $8(b)$]. Therefore, the family fan-out behavior of Δ_K derives from chiral variation of the exchange energy in E_K , not E_{11} .

The dependence of exciton splittings on *d* is currently an open question. One *ab initio* study predicted a 1/*d*² dependence for Δ_K ^{[21](#page-8-15)} and another using the same computational methods employed here has predicted a 1/*d* splitting between zero momentum singlets and triplets. 29 We note that the diameter ranges for these two calculations are nonoverlapping with the former focusing on small diameter tubes and the latter seeking to establish a large diameter scaling limit. Here we find the diameter scaling of Δ_K cannot be uniquely determined from our data set but we are nevertheless able to establish a $1/d^2$ behavior within our calculations for CNTs with diameters lying between 0.6 and 2.5 nm Fig. [9](#page-5-1). This scaling behavior is identical to that of the exchange

FIG. 9. (Color online) Δ_K versus $1/d^2$ calculated by our ETB model. Red (black) squares (circles) are $\nu=1(2)$ CNTs. Lines connect families sharing the same value of $2n+m$ and the green line is a $1/d^2$ fit.

energy responsible for $1/d^2$ behavior in Ref. [21.](#page-8-15) Although the authors of Ref. [29](#page-8-24) did not study Δ_K , the difference in power-law scaling within the exciton manifold predicted by the *same* methods is interesting.

The nature of X_2 remains a controversial topic in the CNT photophysics community. It is not uncommon to see X_2 interpreted as a $\Gamma_{E_{2g}}$ phonon sideband of the bright E_{11} exciton, despite recent theoretical^{12[,13](#page-8-10)} and experimental¹³ works indicating that for certain chiralities the dominant contribution comes from $K_{A_1'}$ phonons. Here we apply our ETB calculation to all CNTs with diameters ranging between 0.6 and 1.33 nm and extract the contributions of the $\Gamma_{E_{2g}}$ and K_{A_1} ¹ phonon modes to the exciton-phonon sideband. The findings

FIG. 10. (Color online) Calculated ratio of $K_{A'_1}$ and $\Gamma_{E_{2g}}$ phonon contributions to X_2 . Red (black) squares (circles) are $\nu = 1(2)$ CNTs. Lines connect families sharing the same value of $2n+m$. The green line labels an average value of 22.8.

FIG. 11. (Color online) Experimental data for (a) $E_{X_2} - E_{11}^{\text{(PLE)}}$ and (b) $E_{X_1} - E_{11}^{\text{(PL)}}$ versus 1/*d*. Lines connect families sharing the same value of $2n+m$ indicated in (a).

strongly support our assignment, as the $K_{A_1'}$ component dominates over that of $\Gamma_{E_{2g}}$ by more than an order of magnitude for all chiralities $[\overline{\text{Fig. 10}}]$ $[\overline{\text{Fig. 10}}]$ $[\overline{\text{Fig. 10}}]$.

The authors of Ref. [20](#page-8-14) recently studied X_1 using PL maps excited near the E_{22} resonance in $(6,5)$, $(7,5)$, and $(10,5)$ CNT samples. Their measurements indicate that E_{X_1} lies ~140 meV below $E_{01}^{(PL)}$ and does not depend on tube diameter or chiral angle.²⁰ Even more recently, the authors of Ref. [11](#page-8-8) observed X_1 in individual CNTs through temperaturedependent PL spectroscopy, reporting E_{X_1} values \sim 130 meV below $E_{11}^{(PL)}$ that exhibit little diameter dependence. In contrast to these studies, we examine both X_1 *and* X_2 , allowing for the observation of chiral behavior in E_{X_1} , E_{X_2} and, most importantly, E_K , which cannot be determined from X_L alone. Katuara plots of our results for E_{X_2}, E_{X_1} relative to $E_{11}^{(PLE)}$ and $E_{11}^{(PL)}$ $E_{11}^{(PL)}$ $E_{11}^{(PL)}$ are summarized in Figs. $11(a)$ and $11(b)$, respectively. Both E_{X_1} and E_{X_2} show similar chiral-family behavior, but the variation in E_{X_1} is approximately 1/5 that found in E_{X_2} and was therefore too small to be resolved by the aforementioned studies. There are two implications of this behavior. First, the sideband spacing approximates two $K_{A'_1}$ phonon energies and depends on CNT chirality. Some dependence of the $K_{A_1'}$ phonon energy on chirality is generically expected from the proximity of the nearest-allowed *k* vectors to the *K*

FIG. 12. (Color online) $E_{X_2} - E_K$ versus 1/*d* for (a) experimental data and (b) our ETB calculation. $E_{X_2} - E_K$ corresponds to half the sideband splitting and nominally gives the energy of the $K_{A_1'}$ phonons. Lines connect families sharing the same value of $2n+m$ indicated in (a).

point, an effect proportional to the phonon dispersion. However, our calculation shows the expected variation due to this effect is much smaller than seen in the data. Figure [12](#page-6-1) shows that ETB predicts $E_{X_2} - E_K(\sim \hbar \omega_K)$ to have a variation \sim 0.5 meV within the chiralities studied whereas our data finds \sim 8 meV. This difference might reflect a limitation of our phonon calculation. Graphene displays Kohn anomalies at the K and Γ points in the iTO and LO graphene phonon branches[.30](#page-8-25)[–33](#page-8-26) Kohn anomalies lead to a strong softening of the phonon energies, resulting in cusps in the dispersion of the optical phonon branch at the Γ and *K* points. Short-range force constant models such as those used here cannot describe this nonanalytic behavior and therefore generically underestimate the phonon dispersion.³² For CNTs, quantization in the circumferential direction significantly enhances the Kohn anomaly in metallic species 34 but in semiconducting CNTs the electronic gap provides a long-distance cutoff to the interatomic force constants and prevents a true Kohn anomaly. Our data suggest the intriguing possibility that there may nevertheless be some enhancement of the iTO phonon dispersion near the *K* point in semiconducting CNTs. A second implication of the data in Fig. [11](#page-6-0) is that the chiral variations in Δ_K mirror those in $\hbar \omega_K$. We have not yet identified a fundamental reason why this should be the case.

V. CONCLUSION

In this manuscript, we present the chirality-dependent behavior of the *K*-momentum dark singlet exciton in 12 highly purified single-chirality CNT samples. These observations complete the characterization of singlet excitons in CNTs. Δ_K exhibits $2n+m$ family behavior, chiral index dependence, and diameter dependence in agreement with ETB calculations and a theoretical study.²¹ This behavior is found to originate from the exchange energy of E_K . We calculate the scaling behavior of Δ_K and suggest a $1/d^2$ law in agreement with Ref. [21.](#page-8-15) The $K_{A_1'}$ phonon energy varies strongly with chirality, in disagreement with our ETB calculations. The origin of this behavior is unclear but may be related to the Kohn anomalies of graphene. Finally, ETB calculations indicate that X_2 arises from $K_{A_1'}$ phonons. We hope that these results can guide the community in further efforts to understand the behavior of dark excitons and their role in the optical properties of carbon nanotubes.

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APPENDIX: INTERACTION KERNEL

We follow the methods outlined in Ref. [24](#page-8-19) to understand the effect of chirality on the interaction kernel²⁴

$$
\mathcal{K}(\mathbf{k}_{c}^{\prime}\mathbf{k}_{v}^{\prime},\mathbf{k}_{c}\mathbf{k}_{v}) = 2\mathcal{K}^{x}(\mathbf{k}_{c}^{\prime}\mathbf{k}_{v}^{\prime},\mathbf{k}_{c}\mathbf{k}_{v}) - \mathcal{K}^{d}(\mathbf{k}_{c}^{\prime}\mathbf{k}_{v}^{\prime},\mathbf{k}_{c}\mathbf{k}_{v}).
$$
\n(A1)

 K^d is the direct interaction, typically responsible for exciton formation, 26 K^x is the exchange interaction, and $\mathbf{k_c}$, $\mathbf{k_v}$ denote wave vectors in the conduction and valence bands. The two interaction terms may be expressed as 24

$$
\mathcal{K}^{x}(\mathbf{k}'_{c}\mathbf{k}'_{v},\mathbf{k}_{c}\mathbf{k}_{v}) = \int d\mathbf{r}' d\mathbf{r} \psi_{\mathbf{k}'_{c}}^{*}(\mathbf{r}') \psi_{\mathbf{k}'_{v}}(\mathbf{r}')
$$

$$
\times v(\mathbf{r}',\mathbf{r}) \psi_{\mathbf{k}_{c}}(\mathbf{r}) \psi_{\mathbf{k}'_{v}}^{*}(\mathbf{r}), \qquad (A2)
$$

$$
\mathcal{K}^{d}(\mathbf{k}_{c}^{\prime}\mathbf{k}_{v}^{\prime},\mathbf{k}_{c}\mathbf{k}_{v}) = \int d\mathbf{r}^{\prime} d\mathbf{r} \psi_{\mathbf{k}_{c}^{\prime}}^{*}(\mathbf{r}^{\prime}) \psi_{\mathbf{k}_{c}}(\mathbf{r}^{\prime})
$$

$$
\times w(\mathbf{r}^{\prime},\mathbf{r}) \psi_{\mathbf{k}_{v}^{\prime}}(\mathbf{r}) \psi_{\mathbf{k}_{v}}^{*}(\mathbf{r}), \qquad (A3)
$$

where *v* and *w* are the bare and screened Coulomb interactions, and ψ are the single quasiparticle wave functions.²⁴

Consider E_{11} created with zero center-of-mass momentum and finite relative momentum near the *K* point, **K**+**k**. The direct term for a momentum transfer **q** in the $q \rightarrow 0$ limit is

$$
\lim_{\mathbf{q}\to\mathbf{0}} \mathcal{K}^d(\mathbf{q}) \propto w(\mathbf{q}\to\mathbf{0}) \int d\mathbf{r}' d\mathbf{r} |\psi_{\mathbf{K}+\mathbf{k},c}(\mathbf{r}')|^2 |\psi_{\mathbf{K}+\mathbf{k},v}(\mathbf{r})|^2.
$$
\n(A4)

This interaction involves the Coulomb potential $w(\mathbf{q})$ for $|{\bf q}|_a \ll 1$, where *a* is the lattice constant and is therefore relatively insensitive to lattice anisotropy.

The exchange interaction

$$
\mathcal{K}^{x}(\mathbf{q}) = v(\mathbf{q}) \int d\mathbf{r}' d\mathbf{r} \psi_{\mathbf{K}+\mathbf{k}+\mathbf{q},c}^{*}(\mathbf{r}') \psi_{\mathbf{K}+\mathbf{k}+\mathbf{q},v}(\mathbf{r}')
$$

$$
\times e^{-i\mathbf{q} \cdot (\mathbf{r}'-\mathbf{r})} \psi_{\mathbf{K}+\mathbf{k},c}(\mathbf{r}) \psi_{\mathbf{K}+\mathbf{k},v}^{*}(\mathbf{r})
$$
(A5)

vanishes in the $q \rightarrow 0$ limit and at finite momentum provides a smaller contribution to the exciton binding energy. The small momentum $|\mathbf{q}|a \ll 1$ terms in the exchange are again insensitive to the lattice anisotropy.

The direct and exchange interactions for E_{11} thus depend only weakly on chiral angle, implying that the observed behavior in Δ_K must originate from E_K . The direct interaction for E_K involves scattering of a particle and hole within their respective valleys, in perfect analogy with the direct interaction for E_{11} . For E_K , the initial and final momenta are \mathbf{k}_c $=$ **K**^{\prime} + **k**, **k**_v $=$ **K** $-$ **k** and **k**_c^{\prime} =**K** \prime + **k**+**q**, **k**_v \prime $=$ **K** $-$ **k**+**q**. The direct interaction in the $q \rightarrow 0$ limit is

$$
\lim_{\mathbf{q}\to\mathbf{0}} \mathcal{K}^d(\mathbf{q}) \propto w(\mathbf{q}\to\mathbf{0}) \int d\mathbf{r}' d\mathbf{r} |\psi_{\mathbf{K'}+\mathbf{k},c}(\mathbf{r}')|^2
$$

$$
\times |\psi_{\mathbf{K}-\mathbf{k},v}(\mathbf{r})|^2. \tag{A6}
$$

The direct interaction for E_K is nearly identical to that of E_{11} , insensitive to chiral angle for the same reasons.

The exchange interaction for E_K is very different from E_{11} since exchanging a particle and hole for an intervalley process requires a minimum momentum transfer of **2K**. All relevant Coulomb-interaction matrix elements will be of order **K** and sensitive to the lattice anisotropy. The exchange may be written as

$$
\mathcal{K}^{x}(\mathbf{q}) = \int d\mathbf{r}' d\mathbf{r} \psi_{\mathbf{K}' + \mathbf{k} + \mathbf{q}, c}^{*}(\mathbf{r}') \psi_{\mathbf{K} - \mathbf{k} + \mathbf{q}, v}(\mathbf{r}')
$$

× $v(\mathbf{q} + 2\mathbf{K}) e^{-i(\mathbf{q} + 2\mathbf{K}) \cdot (\mathbf{r}' - \mathbf{r})} \psi_{\mathbf{K}' + \mathbf{k}, c}(\mathbf{r}) \psi_{\mathbf{K} - \mathbf{k}, v}^{*}(\mathbf{r}).$ (A7)

In the $q \rightarrow 0$ limit, Eq. ([A7](#page-7-0)) becomes

$$
\lim_{\mathbf{q}\to\mathbf{0}} \mathcal{K}^{x}(\mathbf{q}) \propto \int d\mathbf{r}' \psi_{\mathbf{K}'+\mathbf{k},c}^{*}(\mathbf{r}') \psi_{\mathbf{K}-\mathbf{k},v}(\mathbf{r}')
$$
\n
$$
\times \int d\mathbf{r} \psi_{\mathbf{K}'+\mathbf{k},c}(\mathbf{r}) \psi_{\mathbf{K}-\mathbf{k},v}^{*}(\mathbf{r}) v(2\mathbf{K}) e^{-i(2\mathbf{K})\cdot(\mathbf{r}'-\mathbf{r})}. \tag{A8}
$$

Time-reversal symmetry implies that $\psi_{\mathbf{K}'+\mathbf{k},c}(\mathbf{r})$ $=\psi_{K-k,\nu}(r)$ up to a phase factor, allowing Eq. ([A8](#page-7-1)) to be rewritten as

$$
\lim_{\mathbf{q}\to\mathbf{0}} \mathcal{K}^x(\mathbf{q}) \propto \int d\mathbf{r}' |\psi_{\mathbf{K'}+\mathbf{k},c}(\mathbf{r'})|^2 \int d\mathbf{r} |\psi_{\mathbf{K'}+\mathbf{k},c}(\mathbf{r})|^2
$$

$$
\times v(\mathbf{2K}) e^{-i(\mathbf{2K})\cdot(\mathbf{r'}-\mathbf{r})}. \tag{A9}
$$

 $v(2K)$ is responsible for the chiral angle dependence ob-

served in Δ_K . Although $v(2K)$ is very small, the exchange interaction is proportional to the intervalley overlap integrals and produces a much larger anisotropic contribution than the other direct and exchange terms discussed here. Note that **2K** is equivalent to K' where lattice anisotropy effects on the potential are relatively large.

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- 1F. Wang, G. Dukovic, L. E. Brus, and T. F. Heinz, [Science](http://dx.doi.org/10.1126/science.1110265) **308**, 838 ([2005](http://dx.doi.org/10.1126/science.1110265)).
- ²C. L. Kane and E. J. Mele, *[Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.90.207401)* **90**, 207401 (2003).
- ³T. Ando, [J. Phys. Soc. Jpn.](http://dx.doi.org/10.1143/JPSJ.66.1066) **66**, 1066 (1997).
- ⁴H. Zhao and S. Mazumdar, *[Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.93.157402)* **93**, 157402 (2004).
- 5V. Perebeinos, J. Tersoff, and P. Avouris, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.92.257402) **92**, [257402](http://dx.doi.org/10.1103/PhysRevLett.92.257402) (2004).
- 6X. Qiu, M. Freitag, V. Perebeinos, and P. Avouris, [Nano Lett.](http://dx.doi.org/10.1021/nl050227y) **5**, 749 ([2005](http://dx.doi.org/10.1021/nl050227y)).
- 7A. Srivastava, H. Htoon, V. I. Klimov, and J. Kono, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRevLett.101.087402) Lett. **101**[, 087402](http://dx.doi.org/10.1103/PhysRevLett.101.087402) (2008).
- ⁸ J. Shaver, J. Kono, O. Portugall, V. Krstic, G. L. J. A. Rikken, Y. Miyauchi, S. Maruyama, and V. Perebeinos, [Nano Lett.](http://dx.doi.org/10.1021/nl070260f) **7**, 1851 $(2007).$ $(2007).$ $(2007).$
- ⁹ S. Zaric *et al.*, *[Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.96.016406)* **96**, 016406 (2006).
- 10H. Harutyunyan, T. Gokus, A. A. Green, M. C. Hersam, M. Allegrini, and A. Hartschuh, [Nano Lett.](http://dx.doi.org/10.1021/nl9002798) 9, 2010 (2009).
- 11R. Matsunaga, K. Matsuda, and Y. Kanemitsu, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.81.033401) **81**, [033401](http://dx.doi.org/10.1103/PhysRevB.81.033401) (2010).
- 12V. Perebeinos, J. Tersoff, and P. Avouris, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.94.027402) **94**, [027402](http://dx.doi.org/10.1103/PhysRevLett.94.027402) (2005).
- 13O. N. Torrens, M. Zheng, and J. M. Kikkawa, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.101.157401) **101**[, 157401](http://dx.doi.org/10.1103/PhysRevLett.101.157401) (2008).
- ¹⁴ Y. Miyauchi and S. Maruyama, *[Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.74.035415)* **74**, 035415 (2006).
- 15F. Plentz, H. B. Ribeiro, A. Jorio, M. S. Strano, and M. A. Pimenta, *[Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.95.247401)* **95**, 247401 (2005).
- 16S. Berciaud, L. Cognet, P. Poulin, R. B. Weisman, and B. Lounis, [Nano Lett.](http://dx.doi.org/10.1021/nl062933k) 7, 1203 (2007).
- ¹⁷ J. Maultzsch, R. Pomraenke, S. Reich, E. Chang, D. Prezzi, A. Ruini, E. Molinari, M. S. Strano, C. Thomsen, and C. Lienau, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.72.241402) **72**, 241402(R) (2005).
- 18O. Kiowski, K. Arnold, S. Lebedkin, F. Hennrich, and M. M.

Kappes, *[Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.99.237402)* **99**, 237402 (2007).

- 19M. Mohr, J. Maultzsch, E. Dobardzic, S. Reich, I. Milosevic, M. Damnjanovic, A. Bosak, M. Krisch, and C. Thomsen, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRevB.76.035439) **B 76**[, 035439](http://dx.doi.org/10.1103/PhysRevB.76.035439) (2007).
- 20Y. Murakami, B. Lu, S. Kazaoui, N. Minami, T. Okubo, and S. Maruyama, *[Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.79.195407)* **79**, 195407 (2009).
- 21R. B. Capaz, C. D. Spataru, S. Ismail-Beigi, and S. G. Louie, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.74.121401) **74**, 121401(R) (2006).
- $22X$. Tu, S. Manohar, A. Jagota, and M. Zheng, Nature ([London](http://dx.doi.org/10.1038/nature08116)) **460**[, 250](http://dx.doi.org/10.1038/nature08116) (2009).
- 23Z. M. Li, V. N. Popov, and Z. K. Tang, [Solid State Commun.](http://dx.doi.org/10.1016/j.ssc.2004.03.038) **130**[, 657](http://dx.doi.org/10.1016/j.ssc.2004.03.038) (2004).
- ²⁴ J. Jiang, R. Saito, G. G. Samsonidze, A. Jorio, S. G. Chou, G. Dresselhaus, and M. S. Dresselhaus, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.75.035407) **75**, 035407 $(2007).$ $(2007).$ $(2007).$
- 25L. Pietronero, S. Strassler, H. R. Zeller, and M. J. Rice, [Phys.](http://dx.doi.org/10.1103/PhysRevB.22.904) [Rev. B](http://dx.doi.org/10.1103/PhysRevB.22.904) 22, 904 (1980).
- ²⁶M. Rohlfing and S. G. Louie, *[Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.62.4927) 62, 4927 (2000)*.
- 27 C. D. Spataru, S. Ismail-Beigi, L. X. Benedict, and S. G. Louie, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.92.077402) **92**, 077402 (2004).
- 28R. Saito, G. Dresselhaus, and M. S. Dresselhaus, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.61.2981) 61[, 2981](http://dx.doi.org/10.1103/PhysRevB.61.2981) (2000).
- 29V. Perebeinos, J. Tersoff, and P. Avouris, [Nano Lett.](http://dx.doi.org/10.1021/nl051828s) **5**, 2495 $(2005).$ $(2005).$ $(2005).$
- ³⁰ A. C. Ferrari, [Solid State Commun.](http://dx.doi.org/10.1016/j.ssc.2007.03.052) **143**, 47 (2007).
- ³¹ J. Maultzsch, S. Reich, C. Thomsen, H. Requardt, and P. Orde-jon, [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.92.075501) **92**, 075501 (2004).
- 32S. Piscanec, M. Lazzeri, F. Mauri, A. C. Ferrari, and J. Robertson, *[Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.93.185503)* **93**, 185503 (2004).
- 33D. L. Mafra, G. Samsonidze, L. M. Malard, D. C. Elias, J. C. Brant, F. Plentz, E. S. Alves, and M. A. Pimenta, [Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.76.233407) 76[, 233407](http://dx.doi.org/10.1103/PhysRevB.76.233407) (2007).
- 34S. Piscanec, M. Lazzeri, J. Robertson, A. C. Ferrari, and F. Mauri, *[Phys. Rev. B](http://dx.doi.org/10.1103/PhysRevB.75.035427)* 75, 035427 (2007).